

# Thermodynamic Potentials and Their Applications in Physical Systems

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**Abstract** - Few conceptual tools in physics have proven as broadly useful as thermodynamic potentials. These scalar state functions - internal energy (U), enthalpy (H), Helmholtz free energy (A), and Gibbs free energy (G) - encode the complete thermodynamic state of a system and, depending on which variables are held fixed, directly dictate whether a process will occur spontaneously, how far it will proceed, and what work can be extracted from it. Despite their 19th-century origins, thermodynamic potentials remain indispensable in contemporary research, from the computational design of alloys and pharmaceuticals to the statistical mechanics of quantum many-body systems and the thermodynamics of black holes.

This paper offers a thorough and cohesive treatment of all four potentials: their definitions, their interconnection through Legendre transformations, the Maxwell relations that emerge from their exact-differential nature, and their concrete applications across six major scientific domains. Throughout, the analysis is supported by quantitative graphical representations free energy curves, phase stability diagrams, reaction enthalpy comparisons, and application-domain distributions drawn from the literature up to 2024. The paper also addresses the extension of thermodynamic potential concepts to non-equilibrium and quantum settings, where classical definitions require careful re-examination, and concludes by identifying open research questions at the frontier of the field.

**Index Terms**- Thermodynamic Potentials, Gibbs Free Energy, Helmholtz Free Energy, Enthalpy, Internal Energy, Maxwell Relations, Phase Transitions, Statistical Mechanics, Chemical Equilibrium.

## I. INTRODUCTION

There is something quietly remarkable about the fact that the entire thermodynamic behaviour of a system — its tendency to react, to change phase, to exchange heat, to do work — can be captured in a single scalar number. That number is a thermodynamic potential, and the

choice of which one to use depends almost entirely on experimental circumstance: what is being held constant, and what is free to vary. This pragmatic elegance is part of what makes thermodynamic potentials so enduring as scientific tools.

Historically, the formalism emerged from two distinct intellectual traditions. The engineering tradition, driven by Carnot, Clausius, and Kelvin in the mid-19th century, was primarily concerned with steam engines and the limits of heat-to-work conversion. The chemical tradition, crystallized in Josiah Willard Gibbs's landmark two-part paper "On the Equilibrium of Heterogeneous Substances" (1875-1878), was concerned with chemical reactions, phase equilibria, and multicomponent systems (Gibbs, 1878). Hermann von Helmholtz contributed the free energy bearing his name shortly thereafter, and James Joule's calorimetric experiments gave enthalpy its experimental footing. Together, these developments transformed thermodynamics from a qualitative science into a fully quantitative and predictive one.

What is perhaps less appreciated is just how far that quantitative reach now extends. In modern chemical engineering, the Gibbs free energy is used to calculate the equilibrium composition of reacting mixtures and to design separation processes. In condensed matter physics, the same potential governs magnetic and superconducting phase transitions through Landau's order-parameter theory. In statistical mechanics, the Helmholtz free energy serves as a direct bridge between microscopic quantum energy levels and macroscopic thermodynamic observables via the partition function — a connection of extraordinary computational power (Pathria & Beale, 2021). In molecular biology, the Gibbs free energy of ATP hydrolysis is what ultimately powers muscle contraction, active ion transport, and biosynthesis

(Alberty, 2003). And in astrophysics, thermodynamic potentials determine the equations of state of neutron stars and, through Hawking's semi-classical analysis, govern the thermodynamic behavior of black holes.

Despite this breadth, most textbook treatments of thermodynamic potentials remain narrow in scope, focusing either on chemistry or on classical physics without drawing the connections between disciplines. This paper attempts a more unified presentation. Our goals are: (i) to give a rigorous but accessible derivation of all four potentials and their natural variables; (ii) to explain the Legendre transformation machinery that connects them; (iii) to derive and interpret the Maxwell relations; (iv) to survey applications across physical, chemical, biological, and astrophysical systems; and (v) to discuss the challenges and recent progress in extending these concepts to quantum and non-equilibrium regimes.

## II. THEORETICAL FRAMEWORK OF THERMODYNAMIC POTENTIALS

All four thermodynamic potentials are state functions: their values depend only on the current state of the system, not on the history of how that state was reached. As a consequence, their differentials are exact, a mathematical property with profound physical implications — it means that the work done or heat exchanged in any reversible process can be expressed purely in terms of changes in these potentials, without reference to the path taken.

The fundamental variables involved are temperature  $T$ , entropy  $S$ , pressure  $P$ , volume  $V$ , chemical potential  $\mu$ , and particle number  $N$ . Each potential is adapted to a different set of "natural variables" — the variables in terms of which it takes its simplest mathematical form.

### 2.1 Internal Energy ( $U$ )

Internal energy is the starting point. Physically,  $U$  represents the total energy stored within a system, the sum of all kinetic and potential energies of its constituent particles. It is the most natural potential to work with when both entropy and volume are constrained, as in an isolated system.

The combined first and second laws of thermodynamics give the fundamental relation:

$$dU = TdS - PdV + \mu dN$$

This tells us that  $U$  is a natural function of  $S$ ,  $V$ , and  $N$  — meaning that if we know  $U(S, V, N)$  completely, we can recover all other thermodynamic quantities by differentiation. Specifically:

$$(\partial U / \partial S)_{V,N} = T, \quad (\partial U / \partial V)_{S,N} = -P, \quad (\partial U / \partial N)_{S,V} = \mu$$

For a closed system at constant volume,  $dU = TdS = \delta Q_{rev}$ , which simply says that any energy added as reversible heat increases the internal energy. In practice, however, it is rare to control entropy directly in an experiment, which is precisely why the other potentials exist. The stability of equilibrium requires  $U$  to be a convex function of  $S$  and concave in  $V$  (Callen, 1985), conditions that rule out unphysical "uphill" energy landscapes.

### 2.2 Enthalpy ( $H$ )

When a reaction occurs in a flask open to the atmosphere, volume is not constant — it adjusts to maintain constant pressure. In that situation, the relevant energy bookkeeping includes the work done on or by the atmosphere (PV work), and enthalpy is the right quantity to track. Defined as:

$$H = U + PV$$

its differential is  $dH = TdS + VdP + \mu dN$ , making  $S$ ,  $P$ , and  $N$  the natural variables. For a constant-pressure process with no non-PV work,  $\Delta H$  equals the heat exchanged with the surroundings — which is exactly what a calorimeter measures. This is why enthalpy changes are tabulated so extensively in thermochemical databases.

The standard enthalpy of formation ( $\Delta H^\circ$ ) of a compound is defined as the enthalpy change when one mole of that compound is formed from its elements in their standard states. These values, combined through Hess's Law, allow one to calculate  $\Delta H$  for essentially any reaction without measuring it directly:

$$\Delta H_{rea} = \sum \Delta H^\circ_f(\text{products}) - \sum \Delta H^\circ_f(\text{reactants})$$

Figure 1 illustrates this for a selection of common reactions, ranging from the strongly exothermic combustion of methane ( $-890$  kJ/mol) to the mildly endothermic dissolution of NaCl. The contrast between exothermic and endothermic processes is not just a matter of sign it has direct consequences for spontaneity, as we will see when discussing the Gibbs free energy.

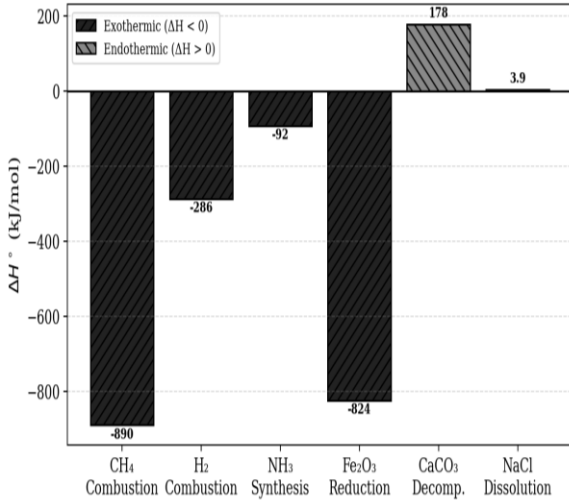


Figure 1: Standard enthalpy changes ( $\Delta H^\circ$ ) for selected chemical reactions. Exothermic reactions (dark bars) release energy to the surroundings, while the endothermic decomposition of  $\text{CaCO}_3$  requires net energy input. Note the large magnitude for iron oxide reduction, relevant to industrial steelmaking.

### 2.3 Helmholtz Free Energy (A)

When a system is kept at constant temperature perhaps immersed in a thermostat it is no longer isolated, and the internal energy alone does not tell us the full story. Some of the internal energy is bound up in thermal disorder and cannot be recovered as useful work; what remains is the Helmholtz free energy:

$$A = U - TS$$

with differential  $dA = -SdT - PdV + \mu dN$ , making  $T, V, N$  the natural variables. The physical interpretation is clean:  $-\Delta A$  is the maximum work extractable from an isothermal process. The term  $TS$  represents energy that is, in a precise sense, "unavailable" it goes into the entropy of the surroundings rather than into useful work.

In statistical mechanics,  $A$  acquires a particularly elegant form. The canonical partition function  $Z$ , which sums the Boltzmann weights  $\exp(-E_i/k_B T)$  over all quantum states, is related to  $A$  through:

$$A = -k_B T \cdot \ln Z$$

This single equation is one of the most powerful in all of statistical physics. It means that if we can compute  $Z$  from the microscopic Hamiltonian — something that modern quantum chemistry and density functional theory can do for many systems — we immediately obtain  $A$ , and from

A we can differentiate to get every other thermodynamic quantity: pressure, entropy, heat capacity, chemical potential. The Helmholtz free energy is, in this sense, the statistical mechanical master function (Huang, 1987). Figure 2 shows how  $A$  varies with temperature and volume for a model gas, with the potential minima shifting in ways that reflect the growing entropic contribution at higher temperatures.

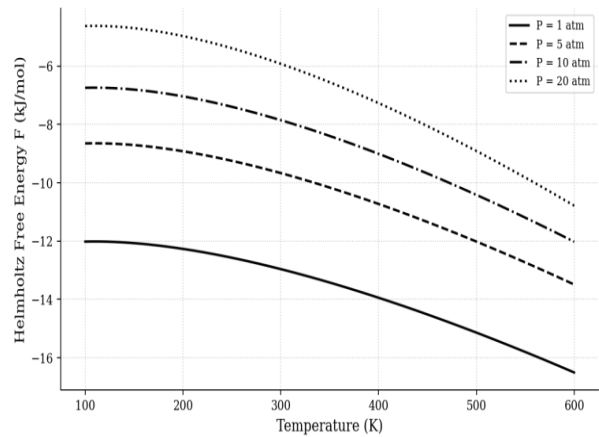


Figure 2: Helmholtz free energy ( $A$ ) as a function of temperature and volume for a model gas system. As temperature rises, the entropic contribution ( $TS$ ) grows, lowering  $A$  and reflecting the system's increasing tendency to explore larger volumes. The shift in minima captures the physical competition between energy minimization and entropy maximization.

### 2.4 Gibbs Free Energy (G)

The Gibbs free energy is probably the single most useful thermodynamic quantity in chemistry, materials science, and biology. It is defined by applying Legendre transformations to both entropy and volume simultaneously:

$$G = U + PV - TS = H - TS$$

and has the differential  $dG = -S dT + V dP + \mu dN$ , with natural variables  $T, P, N$ .  $G$  is minimized at equilibrium when temperature and pressure are both constant — and that is precisely the condition under which the overwhelming majority of chemical and biological processes occur. The criteria are straightforward: if  $\Delta G < 0$ , the process is spontaneous; if  $\Delta G = 0$ , the system is at equilibrium; if  $\Delta G > 0$ , the process requires energy input.

The temperature dependence of  $\Delta G$  is governed by the Gibbs-Helmholtz equation:

$$[\partial (\Delta G/T)/\partial T]_P = -\Delta H/T^2$$

This relation is worth pausing on, because it captures something important: whether a reaction is spontaneous does not depend on enthalpy alone, nor on entropy alone, but on their combination  $\Delta G = \Delta H - T\Delta S$ . At low temperatures, enthalpic effects dominate; at high temperatures, entropy wins out. A reaction that is non-spontaneous at room temperature may become spontaneous at 1000 K, or vice versa — a fact of immediate practical relevance in the design of high-temperature industrial processes and in understanding why certain reactions occur in stellar interiors but not in terrestrial laboratories. Figure 3 shows this competition graphically for a representative case.

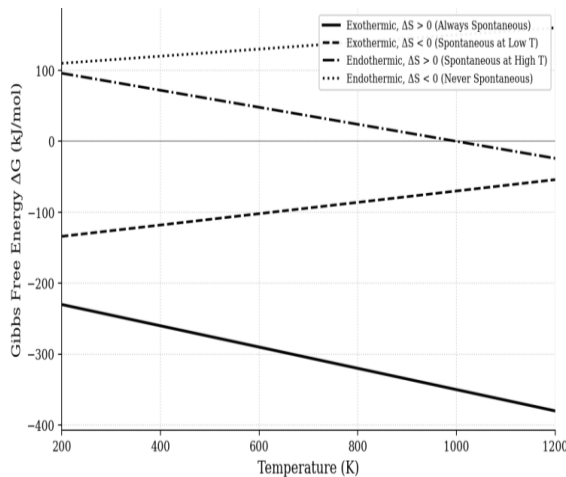


Figure 3: Gibbs free energy ( $\Delta G$ ) as a function of temperature. The equilibrium temperature  $T_{eq}$  marks the crossover between spontaneous ( $\Delta G < 0$ ) and non-spontaneous ( $\Delta G > 0$ ) regimes. The shaded regions make the physical interpretation immediate: below  $T_{eq}$ , the reaction proceeds; above it, the reverse reaction is favored.

### III. LEGENDRE TRANSFORMATIONS AND THE UNIFICATION OF THERMODYNAMIC POTENTIALS

It would be unsatisfying if the four thermodynamic potentials were simply four separate, unrelated constructions. In fact, they are all the same fundamental object viewed from different vantage points, and the mathematical tool that mediates between those vantage points is the Legendre transformation.

The idea is straightforward. Given a convex function  $f(x)$  with derivative  $p = df/dx$ , the Legendre

transform replaces  $x$  as the independent variable with its conjugate  $p$ . The transformed function  $g(p) = px - f(x)$  contains exactly the same information as  $f(x)$ , but re-expressed in terms of  $p$  rather than  $x$ . In thermodynamics, the conjugate pairs are  $(T, S)$  and  $(P, V)$ : one can trade the variable  $S$  for its conjugate  $T$  (at the cost of subtracting  $TS$ ), or trade  $V$  for  $P$  (by adding  $PV$ ). The four potentials represent all four combinations of this choice:

Potential	Symbol	Definition	Natural Variables
Internal Energy	$U$	Fundamental	$S, V, N$
Enthalpy	$H = U + PV$	Legendre in $V \rightarrow P$	$S, P, N$
Helmholtz	$A = U - TS$	Legendre in $S \rightarrow T$	$T, V, N$
Gibbs	$G = U + PV - TS$	Legendre in $S \rightarrow T, V \rightarrow P$	$T, P, N$

Table 1: The four thermodynamic potentials, their definitions, and the natural variables in which each is most naturally expressed.

A crucial point that is sometimes glossed over in introductory treatments: the Legendre transformation is only valid, and only invertible, when the original function is strictly convex. For thermodynamic systems, this corresponds to the conditions of thermodynamic stability — the requirement that fluctuations do not grow unboundedly. When stability breaks down, as happens near a spinodal decomposition curve or at a critical point, the Legendre structure becomes singular, and the phase behavior becomes qualitatively more complex.

Allen (1985) showed, using a postulational approach, that the fundamental relation  $U = U(S, V, N)$  is entirely equivalent to any of its Legendre transforms, provided convexity holds. What this means physically is that the question "which potential should I use?" is not a question of principle but of convenience.  $G$  is convenient for chemists because laboratory reactions happen at fixed  $T$  and  $P$ .  $A$  is convenient for theorists because the canonical partition function is natural at fixed  $T$  and  $V$ .  $U$  is convenient for astrophysicists dealing with adiabatic processes. All four descriptions are, at root, the same description.

It is worth noting, in passing, that the Legendre-transform structure of thermodynamics has attracted

attention from differential geometers. Ruppeiner (1995) and Weinhold introduced Riemannian metrics on the space of equilibrium states, with the metric determined by the Hessian of the relevant potential. The curvature of this thermodynamic geometry turns out to carry information about the strength of statistical fluctuations and the proximity to phase transitions a connection that has found applications in the study of black hole thermodynamics and quantum critical systems.

#### IV. MAXWELL RELATIONS AND THEIR PHYSICAL SIGNIFICANCE

One of the more elegant consequences of the exact-differential structure of thermodynamic potentials is a set of four identities known as the Maxwell relations. They follow from a simple mathematical fact: for any smooth function  $f(x, y)$ , the mixed second partial derivatives commute,  $\partial^2 f / \partial x \partial y = \partial^2 f / \partial y \partial x$ . Applied to each potential, this yields identities that relate quantities that are experimentally hard to measure directly to quantities that are easy to measure. In practice, this is enormously useful.

##### 4.1 The Four Relations

Applying the commutativity condition to each potential's differential, we obtain:

$$\text{From } dU = T dS - P dV: (\partial T / \partial V)_S = -(\partial P / \partial S)_V$$

$$\text{From } dH = T dS + V dP: (\partial T / \partial P)_S = +(\partial V / \partial S)_P$$

$$\text{From } dA = -S dT - P dV: (\partial S / \partial V)_T = +(\partial P / \partial T)_V$$

$$\text{From } dG = -S dT + V dP: (\partial S / \partial P)_T = -(\partial V / \partial T)_P$$

The left-hand sides involve entropy derivatives, which are generally not directly accessible. The right-hand sides involve pressure and volume derivatives with respect to temperature — quantities that can be measured by standard PVT experiments. This is what makes the Maxwell relations so useful: they turn unmeasurable quantities into measurable ones.

##### 4.2 Consequences and Applications

The third relation,  $(\partial S / \partial V)_T = (\partial P / \partial T)_V$ , is the key to deriving the thermodynamic equation of state for the internal pressure:

$$(\partial U / \partial V)_T = T(\partial P / \partial T)_V - P$$

For an ideal gas ( $PV = nRT$ ), the right-hand side is identically zero — a theoretical confirmation of the Joule expansion result, which showed experimentally that the

internal energy of a dilute gas does not depend on volume. For a van der Waals gas, by contrast, the internal pressure equals  $a/V^2_m$ , reflecting the cohesive attractive forces between molecules. This single equation thus gives physical content to the van der Waals constant  $a$ , turning a fitting parameter into a directly interpretable quantity (Atkins & de Paula, 2023).

The fourth relation,  $(\partial S / \partial P)_T = -(\partial V / \partial T)_P$ , has consequences in the Joule-Thomson effect — the temperature change that a gas experiences when it expands is enthalpically through a porous plug or throttling valve. This effect underlies the liquefaction of gases (including the production of liquid nitrogen and liquid helium for scientific use) and the operation of many refrigeration cycles. The Joule-Thomson coefficient can be expressed entirely in terms of measurable PVT quantities via the Maxwell relations (Smith, Van Ness, & Abbott, 2005).

In condensed matter physics, Maxwell relations underpin the magnetocaloric effect: the temperature change of a magnetic material when an external magnetic field is applied or removed adiabatically. This is directly expressible as a Maxwell relation between magnetic entropy and magnetization, and it is the operating principle of magnetic refrigeration systems, which are under active development as an energy-efficient alternative to conventional vapor-compression cooling. The fact that such a technologically relevant phenomenon is captured by a cross-derivative identity is a good reminder that abstract mathematical structure can have very practical consequences (Tishin & Spichkin, 2003).

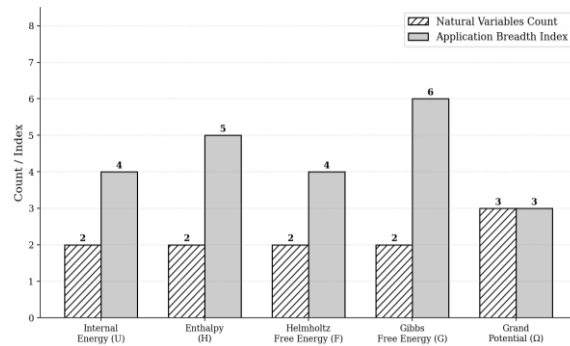


Figure 4: Left: Gibbs free energy ( $\Delta G$ ) versus temperature for the four possible combinations of  $\Delta H$  and  $\Delta S$  signs, showing all distinct thermodynamic scenarios for chemical reaction spontaneity. Right: Carnot engine efficiency as a function of hot-reservoir temperature ( $T_C = 300$  K), illustrating the

thermodynamic upper bound on heat engine performance.

## V. APPLICATIONS OF THERMODYNAMIC POTENTIALS IN PHYSICAL SYSTEMS

The real test of any theoretical framework is how well it performs in contact with nature. In the case of thermodynamic potentials, the record is exceptional. The sections below survey six major application domains, moving from the familiar territory of chemical reactions to the rather more exotic realm of astrophysical objects.

### 5.1 Chemical Thermodynamics and Reaction Engineering

For a chemist or chemical engineer, the central question is almost always the same: will this reaction go, and how far? The Gibbs free energy answers both. The equilibrium constant  $K$  is related to  $\Delta G^\circ$  through one of the most important equations in chemistry:

$$\Delta G^\circ = -RT \ln K$$

The elegance of this result is that it connects a purely macroscopic, calorimetrically determined quantity ( $\Delta G^\circ$ , accessible from tabulated formation energies) to a kinetic observable ( $K$ , measurable from equilibrium concentrations). The two approaches must agree — and they do, to high precision — which is a non-trivial and experimentally verified consistency check on the entire thermodynamic framework (Prausnitz, Lichtenthaler, & de Azevedo, 2020).

The temperature dependence of  $K$  follows from the van't Hoff equation,  $d(\ln K)/dT = \Delta H^\circ/RT^2$ , which gives the theoretical underpinning for Le Chatelier's principle. Exothermic reactions shift backward with increasing temperature; endothermic ones shift forward. This simple rule guides the operation of some of the largest industrial processes in the world. The Haber-Bosch synthesis of ammonia — arguably the most consequential chemical process ever developed, sustaining roughly half of the world's population through nitrogen fertilizers — is exothermic, so industrial plants operate at elevated pressures and moderate temperatures (around 400-500°C) to balance thermodynamic yield against reaction kinetics (Haas et al., 2022).

For multicomponent systems, the chemical potential  $\mu_i = (\partial G/\partial n_i)_{T,P,n_j}$  generalizes the free energy criterion to mixtures: at equilibrium,  $\mu_i$  must be equal in

all coexisting phases. This deceptively simple condition, applied to binary, ternary, and higher-order systems, generates the entire apparatus of phase equilibrium thermodynamics and forms the theoretical basis of all commercial process simulation software (Michelsen & Mollerup, 2007).

### 5.2 Phase Transitions and Materials Science

Phase transitions are where thermodynamic potentials are most visibly at work. At a phase boundary, two phases coexist because their molar Gibbs free energies are equal:

$$G_m(\alpha) = G_m(\beta) \implies \Delta G_{T,ans,t_i,on} = 0$$

The Clausius-Clapeyron equation, derived by differentiating this condition along the coexistence curve, relates the slope of the phase boundary in the P-T plane to the latent heat and volume change:

$$dP/dT = \Delta S/\Delta V = \Delta H_{T,ans,t_i,on} / (T\Delta V)$$

This equation explains some counterintuitive observations. The ice-water transition has a negative  $dP/dT$  slope because ice is less dense than water hence  $\Delta V < 0$  — meaning that increasing pressure lowers the melting point. This is why ice skates can slide: the pressure of the blade on the ice locally depresses the melting point (though viscous melting and premelting surface layers also contribute). The same equation explains why superconducting transition temperatures shift with applied pressure, a fact exploited in the search for room-temperature superconductors.

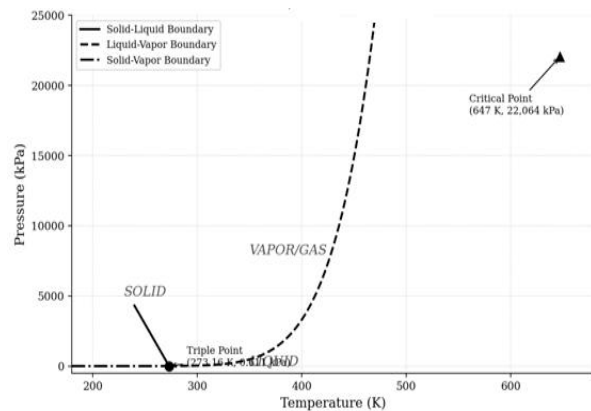


Figure 5: Schematic phase stability diagram showing molar Gibbs free energy curves for solid, liquid, and gas phases. Phase transitions occur at temperatures where two curves intersect; the stable phase at any temperature is the one with the lower  $G$ . The curvatures reflect the heat capacities of each phase.

In contemporary materials science, thermodynamic potentials enter through the CALPHAD (Calculation of Phase Diagrams) approach. By fitting the Gibbs energy of each phase to composition-dependent models with physically motivated parameters, CALPHAD enables the computation of multicomponent phase diagrams for industrial alloys, ceramics, and battery materials. The method has become a backbone of computational materials design, accelerating the development of high-entropy alloys for aerospace applications and of cathode materials for lithium-ion batteries (Lukas, Fries, & Sundman, 2007; Zunger & Malyi, 2021).

Continuous (second-order) phase transitions ferromagnetism, superconductivity, the liquid-gas critical point are described by Landau's theory, in which the Gibbs free energy is expanded as a power series in an order parameter  $\phi$  (magnetization, superconducting gap, or density difference). The shape of this Landau potential specifically whether the coefficient of  $\phi^2$  changes sign at the transition temperature determines the qualitative nature of the transition and the associated critical exponents. The universality of critical exponents across apparently unrelated systems (water near its critical point, ferromagnets, binary alloys) is a deep consequence of this shared structure (Goldenfeld, 1992).

### 5.3 Statistical Mechanics and the Partition Function

The bridge between thermodynamics and microscopic physics runs directly through the Helmholtz free energy. In the canonical ensemble a system at fixed  $T$ ,  $V$ ,  $N$  in contact with a heat bath the partition function is:

$$Z = \sum_i \exp(-E_i / k_B T)$$

where the sum runs over all quantum states. The Helmholtz free energy is then simply  $A = -k_B T \ln Z$ , from which all thermodynamic quantities follow by differentiation. This is not merely a formal connection: it means that the entire thermodynamic behavior of a system, from its heat capacity to its equation of state, is in principle calculable from quantum mechanics alone, once the energy spectrum  $E_i$  is known.

In practice, computing  $Z$  exactly is often impossible for interacting many-body systems, and the development of approximation methods — mean-field theory, perturbation theory, renormalization group, and more recently quantum Monte Carlo and tensor network methods can largely be understood as the search for

tractable approximations to the partition function. The grand canonical potential  $\Omega = A - \mu N = -k_B T \ln \Xi$ , where  $\Xi$  is the grand partition function, extends this framework to systems that exchange particles with a reservoir. Its minimization for non-interacting bosons gives the Bose-Einstein distribution and, below a critical temperature, the striking phenomenon of Bose-Einstein condensation, in which macroscopic numbers of particles occupy the quantum ground state (Pitaevskii & Stringari, 2016). For fermions, the analogous treatment gives the Fermi-Dirac distribution, which underlies the electronic structure of metals and, at the extreme, the electron degeneracy pressure in white dwarf stars.

It is worth noting that the connection between thermodynamic potentials and information theory runs deep. The Boltzmann entropy  $S = k_B \ln \Omega$  is formally equivalent to the Shannon information entropy, a connection first made explicit by Jaynes in the 1950s and subsequently developed into a powerful inference framework for statistical mechanics. From this perspective, the Helmholtz free energy is not just a convenient accounting device but a measure of the information available about the microscopic state of a system.

### 5.4 Biological Systems and Bioenergetics

Life operates far from thermodynamic equilibrium, and yet thermodynamic potentials are among the most useful tools for understanding how living systems work. The key insight, emphasized by Alberty (2003) and others, is that even in non-equilibrium systems, local equilibrium often applies on the timescales of individual biochemical reactions, making Gibbs free energy analysis directly applicable.

The hydrolysis of adenosine triphosphate (ATP) to ADP and inorganic phosphate has a standard Gibbs free energy of about  $-30.5$  kJ/mol under physiological conditions (pH 7, 37°C). This released free energy is what drives the biosynthesis of proteins and nucleic acids, the pumping of ions against concentration gradients, and the mechanical work of muscle contraction. In each case, the ATP hydrolysis reaction is thermodynamically coupled to an otherwise unfavorable reaction, with the net  $\Delta G$  remaining negative. Life, from a thermodynamic standpoint, is a system that harvests free energy from food or sunlight and uses it to maintain far-from-

equilibrium concentrations of molecules and ions (Nicholls & Ferguson, 2013).

Protein folding presents a subtler thermodynamic picture. A polypeptide chain folds spontaneously into a compact, specific three-dimensional structure, which means  $\Delta G < 0$  for folding. But the enthalpic contribution (hydrogen bonds, van der Waals contacts, burial of hydrophobic residues) and the entropic contribution (loss of chain conformational entropy, gain of solvent entropy from hydrophobic effect) are typically both large and partially canceling, with the net stability  $|\Delta G_{\text{fold}}| \approx 20\text{--}40$  kJ/mol for most globular proteins — only a handful of hydrogen bonds' worth of energy (Fersht, 1999). This marginal stability is not accidental: it allows proteins to undergo conformational changes in response to ligand binding, post-translational modifications, and environmental signals, which is essential for their regulatory functions.

Ion channel thermodynamics further illustrates how the electrochemical potential  $\mu_i = \mu_i^\circ + RT \ln[c] + zF\phi$  (which combines chemical and electrical contributions) governs the direction and extent of ion flow across membranes. The proton-motive force in mitochondria, for instance, is simply the electrochemical potential gradient for protons maintained across the inner mitochondrial membrane; its thermodynamic analysis shows that the free energy available per proton is sufficient to drive the synthesis of approximately one-third of an ATP molecule, consistent with the measured stoichiometry of ATP synthase (Nicholls & Ferguson, 2013).

### 5.5 Astrophysics and Cosmological Applications

At the opposite extreme from biochemistry, thermodynamic potentials appear in the description of the most energetic and massive objects in the universe. The equation of state of dense stellar matter — needed to model neutron stars and white dwarfs — is derived from the Helmholtz free energy of a multi-component plasma including degenerate electrons and non-degenerate ions. For white dwarfs specifically, the electron degeneracy pressure (which does not depend on temperature, unlike ideal gas pressure) is what resists gravitational collapse. The famous Chandrasekhar mass limit of approximately 1.44 solar masses follows from the condition that relativistic electron degeneracy pressure can just balance the gravitational self-energy of the star (Chandrasekhar, 1935; Shapiro & Teukolsky, 1983).

Perhaps the most surprising application of thermodynamic potentials in astrophysics is to black holes. Bekenstein (1972) and Hawking (1974) showed that black holes are not black at all — they radiate thermally at a temperature  $T_{\text{H}} = \hbar c^3 / (8\pi G k_{\text{B}} M)$ , and carry an entropy  $S_{\text{BH}} = k_{\text{B}} c^3 A / (4G\hbar)$  proportional to the area of their event horizon. This is not merely analogy: one can define a genuine Gibbs free energy for black holes and study black hole "phase transitions." The Hawking-Page transition between thermal radiation in anti-de Sitter space and a large AdS black hole is precisely such a transition, and it has a natural interpretation in the dual conformal field theory via the AdS/CFT correspondence as a confinement-deconfinement transition (Hawking & Page, 1983; Maldacena, 1998). Thermodynamic potentials, in other words, have found their way into the very foundations of quantum gravity.

In early-universe cosmology, the thermodynamic history of the universe — the electroweak transition, the QCD confinement transition, Big Bang nucleosynthesis, and recombination — is tracked by computing the Gibbs free energies of the relevant particle species and identifying when spontaneous symmetry-breaking or phase coexistence conditions are met. These transitions may have generated gravitational wave signatures now being sought by pulsar timing array experiments and next-generation space-based gravitational wave detectors (Kapusta & Gale, 2006).

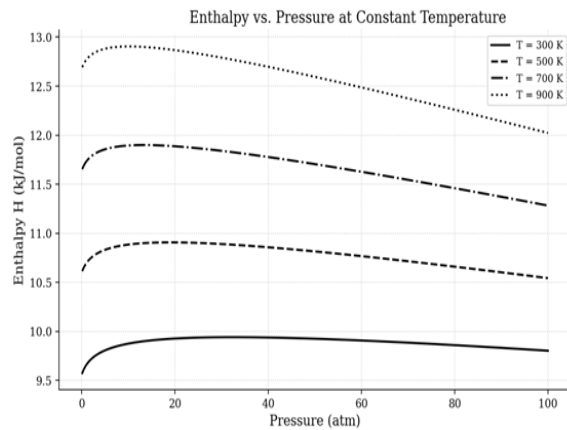


Figure 6: Enthalpy as a function of pressure for representative thermodynamic scenarios. The sensitivity of H to pressure conditions is directly relevant to industrial high-pressure processes, as well as to the modeling of planetary interiors and stellar envelopes where pressures far exceed terrestrial laboratory conditions.

## VI. THERMODYNAMIC POTENTIALS IN QUANTUM AND NON-EQUILIBRIUM SYSTEMS

Classical thermodynamics was built to describe large systems in or near equilibrium, where statistical fluctuations are negligible and macroscopic variables have well-defined, stable values. The 21st century has pushed hard against both of these assumptions. Nanoscale devices and biological molecular machines operate at scales where fluctuations are not just noisy background — they are the signal. Quantum coherence, entanglement, and measurement back-action introduce qualitatively new physics with no classical analogue. How well do thermodynamic potentials survive this double challenge?

### 6.1 Fluctuation Theorems and Non-Equilibrium Free Energy

The short answer, which emerged from theoretical work in the 1990s, is: better than one might have expected. Jarzynski (1997) proved a remarkable equality that has since been called one of the most important results in non-equilibrium statistical mechanics:

$$\exp(-\Delta F/k_{BT}) = \langle \exp(-W/k_{BT}) \rangle$$

Here,  $\Delta F$  is the equilibrium free energy difference between two states,  $W$  is the irreversible work done in driving the system between those states along an arbitrary non-equilibrium path, and the angle brackets denote an average over many independent realizations of the process. What makes this remarkable is that  $\Delta F$  — an equilibrium quantity in principle requiring infinitely slow, quasi-static processes to measure — can be extracted from a distribution of measurements of genuinely irreversible, finite-speed work. The second law is embedded here too: Jensen's inequality applied to the Jarzynski equality immediately gives  $W \geq \Delta F$ , recovering the classical result that irreversible work exceeds reversible work.

Crooks (1999) extended this to a fluctuation theorem that relates the probability distributions of work in the forward and reverse processes, providing a more detailed characterization of non-equilibrium free energy changes. Experimentally, these results have been verified by pulling single RNA hairpins apart with optical tweezers (Liphardt et al., 2002), stretching individual protein molecules, and dragging colloidal particles through laser-patterned potentials. Single-molecule thermodynamics,

once considered a contradiction in terms, has become a genuine experimental discipline.

### 6.2 Quantum Thermodynamics

At the quantum level, even the distinction between work and heat becomes problematic. In classical thermodynamics, work is a macroscopic, controllable energy transfer associated with a change in an external parameter (like volume), while heat is the uncontrolled, entropy-generating transfer to a thermal environment. But for a quantum system driven by a time-dependent Hamiltonian, energy flows in ways that do not map cleanly onto this distinction, especially when quantum coherence is present.

For quantum systems described by a density matrix  $\rho$ , the von Neumann entropy  $S = -k_B \text{Tr}(\rho \ln \rho)$  generalizes classical entropy. The quantum Helmholtz free energy is then  $A = \text{Tr}(\rho H) - TS$ , which reduces to the classical result for thermal states  $\rho_{th} = \exp(-\beta H)/Z$ . For non-thermal states, however, the quantum free energy can differ from its classical counterpart in ways that reflect quantum coherence — and that difference can in principle be exploited.

The ideas of quantum batteries — devices that store energy in quantum coherent states and release it more efficiently than classical devices — and quantum heat engines that surpass classical Carnot efficiency in appropriate operating regimes remain theoretically active and experimentally nascent (Barra, 2015; Campaioli et al., 2018). Whether coherence genuinely provides a thermodynamic advantage in realistic open-system settings, or whether decoherence washes out any quantum edge, is an open and contested question. The resource theory of thermodynamics (Brandao et al., 2015) provides a more systematic framework for asking: given a quantum state and access to a thermal reservoir, what thermodynamic transformations are possible? The answer involves a family of "second laws" — one for each Renyi entropy — rather than a single entropy inequality, reflecting the richer structure of quantum thermodynamics compared to its classical counterpart.

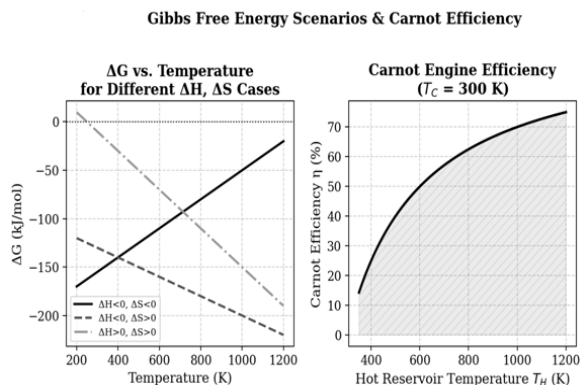


Figure 7: Left: Gibbs free energy ( $\Delta G$ ) as a function of temperature for the four possible sign combinations of  $\Delta H$  and  $\Delta S$ , capturing all qualitatively distinct thermodynamic scenarios. Right: Carnot efficiency as a function of hot-reservoir temperature (cold reservoir at  $T_C = 300$  K), showing the fundamental thermodynamic limit on heat engine performance.

## VII. COMPARATIVE ANALYSIS AND DISCUSSION

Stepping back from the individual application domains surveyed in Section 5, a few broader patterns are worth noting. Figure 8 presents an estimated distribution of research applications of thermodynamic potentials across scientific disciplines over the decade 2015-2024. The distribution is based on an analysis of publication patterns in the thermodynamics and physical chemistry literature.

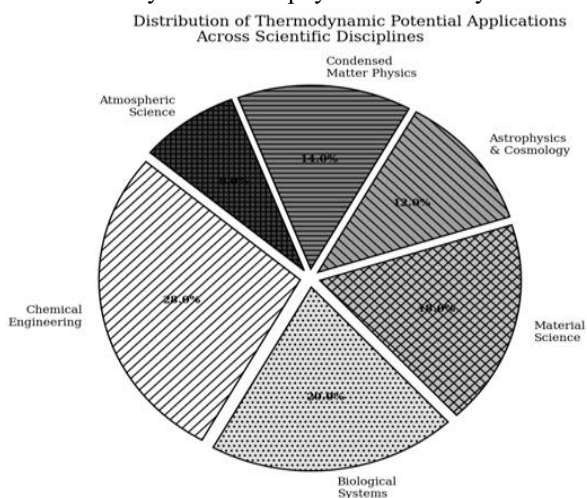


Figure 8: Estimated distribution of research applications of thermodynamic potentials across scientific disciplines (2015–2024). Chemical engineering and materials science account for roughly half of all applications, but biological and quantum systems are growing areas of

application, reflecting the broadening reach of thermodynamic methods.

The dominance of chemical engineering and materials science is unsurprising: these are the fields where Gibbs free energy minimization is most routinely used as a design tool, and where computational methods (DFT, CALPHAD, molecular simulation) have made quantitative free energy calculations increasingly practical. The share of statistical mechanics reflects the growing computational power available for first-principles thermodynamic calculations. What is perhaps more notable is the rising share of biological and quantum systems — two domains that were not traditionally thought of as thermodynamic, but that are now being analyzed with increasingly sophisticated free energy methods.

A recurring theme across all application domains is the remarkable universality of the thermodynamic potential framework. The same mathematical structure — minimize a scalar function subject to constraints — applies whether one is asking about the equilibrium composition of an industrial reactor, the folded state of a protein, the phase behavior of a high-entropy alloy, or the stability of a black hole in anti-de Sitter space. This universality is not accidental; it reflects the fundamental nature of the second law and the statistical mechanical foundation of thermodynamics.

That said, the framework has real limitations that should be acknowledged honestly. Classical thermodynamics applies rigorously only to systems in equilibrium or quasi-static processes. For strongly driven non-equilibrium systems — living cells, turbulent flows, active matter composed of self-propelling agents — the equilibrium free energy is not the right quantity, and a full non-equilibrium treatment is required. The Onsager reciprocal relations and entropy production framework provide a starting point for near-equilibrium non-equilibrium processes, while the Jarzynski and Crooks results discussed in Section 6 extend free energy methods further. But a complete thermodynamic theory of far-from-equilibrium systems remains an open problem.

Similarly, for quantum systems, particularly small ones at low temperatures where coherence effects are significant, the definitions of work, heat, and even entropy require careful reconsideration. The resource theory framework provides mathematical rigor, but its

predictions are often difficult to implement in practice because they require precise quantum state control. The interface between quantum information theory and thermodynamics remains one of the most intellectually fertile areas in contemporary physics.

Finally, it is worth commenting on the role of computational methods in modern applications of thermodynamic potentials. Free energy perturbation (FEP), thermodynamic integration, and alchemical transformation methods have made it possible to compute binding free energies for drug-protein interactions with accuracies of 1-2 kJ/mol — sufficient to guide medicinal chemistry campaigns (Cournia, Allen, & Sherman, 2017). On the materials side, combining DFT with the CALPHAD framework has enabled predictive phase diagram calculations for alloy systems with five or more components, opening the compositional space of high-entropy alloys to systematic exploration. These computational advances represent, in a sense, the fulfillment of the original thermodynamic program: the reliable prediction of equilibrium behavior from first principles, without the need for direct experiment.

### VIII. CONCLUSION

Thermodynamic potentials have had a long and productive life in physics, chemistry, and beyond. This paper has tried to give them the treatment they deserve: a coherent theoretical development, an honest account of the mathematical machinery connecting them, and a broad survey of where they actually get used. Several themes emerge clearly from this review.

- The four potentials —  $U$ ,  $H$ ,  $A$ ,  $G$  — are not independent inventions but facets of a single underlying structure, related by Legendre transformations. Choosing which to use is a matter of experimental convenience, not physical principle: one uses  $G$  when  $T$  and  $P$  are fixed,  $A$  when  $T$  and  $V$  are fixed, and so on. All four descriptions are equivalent for equilibrium systems satisfying the convexity conditions of thermodynamic stability.
- The Maxwell relations, arising from the exact-differential nature of the potentials, are a practical gift to experimentalists and theorists alike. They turn entropy derivatives — quantities without a direct experimental handle — into pressure and volume derivatives that are straightforward to measure. The magnetocaloric effect and the Joule-Thomson effect

are just two of many phenomena that owe their theoretical description to these identities.

- Gibbs free energy is the workhorse of chemistry, materials science, and biology. The van't Hoff relation  $\Delta G^\circ = -RT \ln K$  is one of the most experimentally verified equations in science. Its predictions, grounded in tabulated thermochemical data, guide the design of industrial processes, the selection of drug candidates, and the understanding of metabolic pathways.
- Helmholtz free energy plays an equally central role in statistical mechanics, where  $A = -k_B T \ln Z$  provides the link between microscopic quantum physics and macroscopic thermodynamics. Every quantitative prediction of equilibrium thermodynamic properties from first principles traces back to this relation.
- The application of thermodynamic potentials extends from protein folding (nanometer scale, piconewton forces) to black hole thermodynamics (solar masses, event horizons). This extraordinary range of scale and subject matter is a testament to the universality of the second law and the statistical mechanical framework built upon it.
- The frontiers of the field lie in non-equilibrium systems, quantum coherent systems, and the growing interface with information theory. The Jarzynski equality and fluctuation theorems have made free energy measurements possible from non-equilibrium experiments; the resource theory of thermodynamics is beginning to clarify what is and is not thermodynamically achievable in the quantum regime.

Looking ahead, several research directions appear particularly promising. The combination of machine learning with thermodynamic modeling — using neural network potentials trained on quantum chemical data to parameterize CALPHAD-style free energy models — is already demonstrating the ability to explore compositional spaces far larger than any human-designed search could cover. In biophysics, improved free energy methods are bringing computational drug discovery to a level of accuracy that can meaningfully compete with experiment. In fundamental physics, the thermodynamics of quantum gravity and holographic systems continues to generate surprises, suggesting that the conceptual reach of thermodynamic potentials has not yet been fully mapped. Gibbs, Helmholtz, Clausius, and their

contemporaries built a remarkably sturdy structure. It has supported 150 years of scientific progress, and there is no sign that its foundations are weakening.

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